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QCD analysis of xF_3 structure functions in deep-inelastic scattering: Mellin transform by Gegenbauer polynomial up to N³LO approximation

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This paper provides a thorough examination of the xF_3 structure functions in deep-inelastic scattering through a comprehensive QCD analysis. Our approach harnesses sophisticated mathematical techniques, namely the Mellin transform combined with Gegenbauer polynomials. We have employed the Jacobi polynomials approach for analysis, conducting investigations at three levels of precision: Next-to-Leading Order (NLO), Next-to-Next-to-Leading Order (N²LO), and Next-Next-Next-to-Leading Order (N³LO). We have performed a comparison of our sets of valence-quark parton distribution functions with those of recent research groups, specifically CT18 and MSHT20 at NLO and N²LO, and MSTH23 at N³LO, which are concurrent with our current analysis. The combination of Mellin transforms with Gegenbauer polynomials proves to be a powerful tool for investigating the xF_3 structure functions in deep-inelastic scattering and the results obtained from our analysis demonstrate a favorable alignment with experimental data.

I. INTRODUCTION

Quantum Chromodynamics (QCD) is the fundamental theory of strong interactions, describing the behavior of quarks and gluons, the building blocks of protons, neutrons, and other hadrons. Deep-Inelastic Scattering (DIS) experiments have been a cornerstone in the study of QCD, providing crucial insights into the internal structure of nucleons and the distribution of quarks and gluons within them. Among the various observables in DIS, the xF_3 structure functions hold particular significance as they encode essential information about the parton distribution functions (PDFs) within the nucleon.

The xF_3 structure functions, related to the charged current DIS, play a pivotal role in testing the predictions of QCD and probing the dynamics of quarks and gluons at different energy scales. The study of xF_3 [1–21] has witnessed significant progress over the years, with advances in both experimental techniques and theoretical frameworks. To extract precise information from experimental data and interpret it in the context of QCD, sophisticated theoretical tools and mathematical techniques are required.

In this context, the Mellin transform [24], which allows for the determination of moments of the proton structure functions, has emerged as an effective tool for the quantum chromodynamics (QCD) analysis of the xF_3 structure functions. The Mellin transform facilitates the improvement of approximations, leading to more accurate predictions for scaling violations and the evolution of parton distribution functions (PDFs) with changing momentum scales. We have used Jacobi polynomials to perform the transformation of the evolved functions from the Mellin space to the Bjorken x space, which is an important step in our analysis, as it allows us to compare our results with experimental data. Additionally, we leverage Gegenbauer polynomials for PDFs parameterization, which offer orthogonality and flexibility in function approximation. We show that using the Gegenbauer polynomial expansion method and the next-to-next-tonext-to-leading order (NNNLO) approximation, which are novel and accurate techniques in this context.

Using Gegenbauer polynomials in models can increase the precision and quality of the approximations and the theoretical models. The best values for the parameters of these polynomials can be chosen by fitting them to experimental data.

This paper presents a comprehensive QCD analysis of the proton structure function xF_3 using neutrino-nucleus scattering data from CCFR [25], NuTeV [26] and CHO-RUS [27] experiments. We explain the theoretical framework of QCD and the Jacobi polynomial approach, and the key concepts for the subsequent analysis.

We conclude that the QCD analysis of the xF_3 structure function using Mellin transforms with Jacobi polynomials and Gegenbauer polynomials for PDFs parameterization. This methods is a fast, precise, and direct way to calculate the final structure function with high accuracy. It allows us to extract the valence-quark distribu-

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tion functions from the neutrino-nucleus scattering data without complicated calculations in the kinematic region of interest.

This article is organized as follows. In Sec. II, we introduce the PDFs parametrization and the theoretical framework for the Mellin transform. In Sec. III, we describe the Jacobi polynomials and the QCD fit procedure. In Sec. IV, we present the fit results. In Sec. V, we discuss the GLS sum rule. The final section, Sec. VI, summarizes and concludes the article.

II. THEORETICAL FRAMEWORK TO MELLIN TRANSFORM

In charged-current neutrino deep inelastic scattering (DIS) processes, a neutrino ν ($\bar{\nu}$) interacts with a quark inside the nucleon through the exchange of a virtual W^{\pm} boson. The nonsinglet structure function $xF_3(x,Q^2)$, which arises from the parity-violating weak interaction, characterizes the momentum density of partons, including both valence quarks and antiquarks, within the nucleon.

In the quark-parton model (QPM), the structure functions $xF_3^{\nu p}$ and $xF_3^{\bar{\nu}p}$ for neutrino-proton and antineutrino-proton interactions are given by changing the signs of the antiquark distributions in the expressions for $F_2^{\nu p}$ and $F_2^{\bar{\nu}p}$. By considering $F_2 = 2xF_1$, one can have the above $F_2^{\nu p}$ and $F_2^{\bar{\nu}p}$ structure functions in terms of PDFs, $F_2^{\nu p} = 2x(d + s + \bar{u} + \bar{c})$, and $F_2^{\bar{\nu}p} = 2x(u + c + \bar{d} + \bar{s})$. By changing the signs of $\bar{u}, \bar{d}, \bar{s}$, and \bar{c} , The correct expressions for $xF_3^{\nu p}$ and $xF_3^{\bar{\nu}p}$ are:

$$xF_{3}^{\nu p} = 2x(d+s-\bar{u}-\bar{c}) , xF_{3}^{\bar{\nu}p} = 2x(u+c-\bar{d}-\bar{s}) .$$
 (1)

By considering $u \equiv u_v + \bar{u}$ and $d \equiv d_v + \bar{d}$ and combining the above equations, the structure function xF_3 is as follows:

$$xF_3^{(\nu+\bar{\nu})p} = xF_3^{\nu p} + xF_3^{\bar{\nu}p} = 2x(u_v + d_v) + 2x(s-\bar{s}) + 2x(c-\bar{c})$$
(2)

So, one can have the average of the neutrino and antineutrino nucleon structure function as follows:

$$xF_{3}^{N}(x,Q^{2}) = \frac{1}{2} \left(xF_{3}^{\nu N} + xF_{3}^{\bar{\nu}N} \right) (x,Q^{2}) = \frac{1}{2} \left([xF_{3}^{(\nu+\bar{\nu})p} + xF_{3}^{(\nu+\bar{\nu})n}]/2 \right) (x,Q^{2}) .$$
(3)

However, due to the isospin symmetry, $xF_3^{(\nu+\bar{\nu})p} = xF_3^{(\nu+\bar{\nu})n}$, the average of the neutrino and antineutrino nucleon structure is

$$xF_3^N(x,Q^2) = \frac{1}{2} xF_3^{(\nu+\bar{\nu})p}(x,Q^2)$$

= $[x(u_v+d_v)+x(s-\bar{s})+x(c-\bar{c})](x,Q^2)$.
(4)

It is important to recognize that the differences between the strange quark and its antiquark $s - \bar{s}$, as well as the charm quark and its antiquark $c - \bar{c}$, are typically negligible. Consequently, the average structure of the nucleon as probed by neutrinos and antineutrinos predominantly reflects the distribution of valence quarks

$$xF_3^N(x,Q^2) = (xu_v + xd_v)(x,Q^2) .$$
 (5)

where the combinations $d_v \equiv d - \bar{d}$ and $u_v \equiv u - \bar{u}$ correspond to the valence densities of down and up quarks, respectively, in the proton. The quantities s(x) and c(x) represent the distributions of strange and charm quarks, while $\bar{c}(x)$ is the distribution of charm antiquarks.

When experimental data is reported by collaborations such as CCFR [25], NuTeV data [26] and CHORUS data [27].

This expression characterizes the momentum density of partons, including valence quarks and antiquarks, within the nucleon and provides essential insights into the quarkgluon dynamics in the proton.

For the current analysis, we employ the following standard parameterizations for the valence distributions, xu_v and xd_v , using Gegenbauer polynomials:

$$xu_{v} = \mathcal{N}_{u} x^{\alpha_{u_{v}}} (1-x)^{\beta_{u_{v}}} \left(1 + \sum_{i=1}^{3} a_{u}^{i} C^{\frac{7}{2}}(i, 1-2x)\right),$$
(6)

$$xd_{v} = \frac{\mathcal{N}_{d}}{\mathcal{N}_{u}}(1-x)^{\beta_{d_{v}}}xu_{v}.$$

$$xd_{v} = \mathcal{N}_{d}x^{\alpha_{u_{v}}}(1-x)^{\beta_{u_{v}}+\beta_{d_{v}}}\left(1+\sum_{i=1}^{3}a_{u}^{i}C^{\frac{7}{2}}(i,1-2x)\right),$$
(7)

where $Q_0^2 = 1 GeV^2$ is the input scale and the $C^{\frac{7}{2}}(i, 1-2x)$ are Gegenbauer polynomials. The normalizations \mathcal{N}_u and \mathcal{N}_d are being fixed by $\int_0^1 u_v dx = 2$ and $\int_0^1 d_v dx = 1$, respectively.

We use the neutrino-nucleus data so we take into account the nuclear effects and use the nuclear weight function in the structure function xF_3 calculations. We select the new form for xu_v and xd_v that follows the following expression

$$xu_v^A = \mathcal{W}_{u_v} \frac{Zxu_v + Nxd_v}{A}$$
$$xd_v^A = \mathcal{W}_{d_v} \frac{Nxu_v + Zxd_v}{A}$$
(8)

The shape of the weight function is as follows:

$$\mathcal{W}_{u_v} = 1 + \left(1 - \frac{1}{A^{1/3}}\right) \frac{A_u + c_1 x + c_2 x^2 + c_3 x^3}{(1 - x)^{0.4}}$$
$$\mathcal{W}_{d_v} = 1 + \left(1 - \frac{1}{A^{1/3}}\right) \frac{A_d + c_1 x + c_2 x^2 + c_3 x^3}{(1 - x)^{0.4}}$$
(9)



Figure 1: The nuclear weight function calculated in refs. [22, 23] in NLO and NNLO approximation for ${}^{56}Fe$.

This distribution of nuclear weight was calculated in the refs. [22, 23] in the NLO and NNLO approximation and are shown in Figs. 1 and 2. In our analysis, we used the NNLO weight function for the NNNLO approximation, because no group has obtained a weight function in this approximation, i.e., NNNLO, for the nuclear structure functions. The deep inelastic scattering data of the neutrino nucleus are for iron and lead, and we selected A=56, Z=26 and N=A-Z to represent the number of neutrons for iron and A=208, Z=82 for lead.

The Mellin transform of these functions are defined as

$$u_v^A(N,Q_0^2) = \int_0^1 x u_v^A(x,Q_0^2) x^{n-2} dx$$

$$d_v^A(N,Q_0^2) = \int_0^1 x d_v^A(x,Q_0^2) x^{n-2} dx.$$
(10)

The evolution equation of the nonsinglet structure function $xF_3(x, Q^2)$ in Mellin space, extended to the NLOloop order, can be found in the reference [24].

$$F_{3}(N,Q^{2}) = \left(1 + a C_{3}^{(1)}(N)\right) \times F_{3}(N,Q_{0}^{2}) \left(\frac{a}{a_{0}}\right)^{-\hat{P}_{0}(N)/\beta_{0}} \left\{1 - \frac{1}{\beta_{0}}(a - a_{0}) \left[\hat{P}_{1}^{+}(N) - \frac{\beta_{1}}{\beta_{0}}\hat{P}_{0}(N)\right]\right\}$$
(11)

The evolution equation of $xF_3(x, Q^2)$, extended to the N²LO-loop order, can be found in the reference [28, 29].



Figure 2: The nuclear weight function calculated in the refs. [22, 23] in NLO and NNLO approximation for ²⁰⁸*Pb*.

Within this framework, the nonsinglet structure functions can be expressed as follows:

$$F_{3}(N,Q^{2}) = \left(1 + a C_{3}^{(1)}(N) + a^{2} C_{3}^{(2)}(N)\right) \\ \times F_{3}(N,Q_{0}^{2}) \left(\frac{a}{a_{0}}\right)^{-\hat{P}_{0}(N)/\beta_{0}} \\ \left\{1 - \frac{1}{\beta_{0}}(a - a_{0}) \left[\hat{P}_{1}^{+}(N) - \frac{\beta_{1}}{\beta_{0}}\hat{P}_{0}(N)\right] \\ - \frac{1}{2\beta_{0}} \left(a^{2} - a_{0}^{2}\right) \left[\hat{P}_{2}^{+}(N) - \frac{\beta_{1}}{\beta_{0}}\hat{P}_{1}^{+}(N) \\ + \left(\frac{\beta_{1}^{2}}{\beta_{0}^{2}} - \frac{\beta_{2}}{\beta_{0}}\right)\hat{P}_{0}(N)\right] \\ + \frac{1}{2\beta_{0}^{2}} \left(a - a_{0}\right)^{2} \left(\hat{P}_{1}^{+}(N) - \frac{\beta_{1}}{\beta_{0}}\hat{P}_{0}(N)\right)^{2}\right\}.$$
(12)

The NLO and N²LO Wilson coefficient functions $C_3^{(1)}$ and $C_3^{(2)}$ in Mellin N space can be determined easily using the references [30, 31]. The splitting functions in Mellin N space can be found in Refs. [32–36].

The expansion coefficients β_k of the β -function of QCD are known up to k = 3, corresponding to the N³LO (nextto-next-to-next-to-leading order) [37–39]. These β_k coefficients are important for determining the evolution of the strong coupling constant α_s with the scale Q^2 and are significant in the perturbative calculations of various

QCD processes:

$$\begin{aligned} \beta_0 &= 11 - 2/3 \ n_f \ , \\ \beta_1 &= 102 - 38/3 \ n_f \ , \\ \beta_2 &= 2857/2 - 5033/18 \ n_f + 325/54 \ n_f^2 \ , \\ \beta_3 &= 29243.0 - \ 6946.30 \ n_f + 405.089 \ n_f^2 + 1093/729 \ n_f^3 \ , \end{aligned}$$
(13)

here n_f stands for the number of effectively massless quark flavors and β_k denote the coefficients of the usual four-dimensional \overline{MS} beta function of QCD.

Within this framework, the nonsinglet structure functions can be expressed as follows in N^3LO :

$$F_{3}(N,Q^{2}) = \left(1 + a_{s} C_{3,NS}^{(1)}(N) + a_{s}^{2} C_{3,NS}^{(2)}(N) + a_{s}^{3} C_{3,NS}^{(3)}(N)\right) F_{3}(N,Q_{0}^{2}) \\ \times \left(\frac{a_{s}}{a_{0}}\right)^{-\hat{P}_{0}(N)/\beta_{0}} \left\{1 - \frac{1}{\beta_{0}}(a_{s} - a_{0}) \left[\hat{P}_{1}^{+}(N) - \frac{\beta_{1}}{\beta_{0}}\hat{P}_{0}(N)\right] \\ - \frac{1}{2\beta_{0}}\left(a_{s}^{2} - a_{0}^{2}\right) \left[\hat{P}_{2}^{+}(N) - \frac{\beta_{1}}{\beta_{0}}\hat{P}_{1}^{+}(N) + \left(\frac{\beta_{1}^{2}}{\beta_{0}^{2}} - \frac{\beta_{2}}{\beta_{0}}\right)\hat{P}_{0}(N)\right] \\ + \frac{1}{2\beta_{0}^{2}}\left(a_{s} - a_{0}\right)^{2} \left(\hat{P}_{1}^{+}(N) - \frac{\beta_{1}}{\beta_{0}}\hat{P}_{0}(N)\right)^{2} \\ - \frac{1}{3\beta_{0}}\left(a_{s}^{3} - a_{0}^{3}\right) \left[\hat{P}_{3}^{+}(N) - \frac{\beta_{1}}{\beta_{0}}\hat{P}_{2}^{+}(N) + \left(\frac{\beta_{1}^{2}}{\beta_{0}^{2}} - \frac{\beta_{2}}{\beta_{0}}\right)\hat{P}_{1}^{+}(N) \\ + \left(\frac{\beta_{1}^{3}}{\beta_{0}^{3}} - 2\frac{\beta_{1}\beta_{2}}{\beta_{0}^{2}} + \frac{\beta_{3}}{\beta_{0}}\right)\hat{P}_{0}(N)\right] \\ + \frac{1}{2\beta_{0}^{2}}\left(a_{s} - a_{0}\right)\left(a_{0}^{2} - a_{s}^{2}\right)\left(\hat{P}_{1}^{+}(N) - \frac{\beta_{1}}{\beta_{0}}\hat{P}_{0}(N)\right) \\ \times \left[\hat{P}_{2}(N) - \frac{\beta_{1}}{\beta_{0}}\hat{P}_{1}(N) - \left(\frac{\beta_{1}^{2}}{\beta_{0}^{2}} - \frac{\beta_{2}}{\beta_{0}}\right)\hat{P}_{0}(N)\right] \\ - \frac{1}{6\beta_{0}^{3}}\left(a_{s} - a_{0}\right)^{3}\left(\hat{P}_{1}^{+}(N) - \frac{\beta_{1}}{\beta_{0}}\hat{P}_{0}(N)\right)^{3}\right\}.$$
(14)

and

$$F_3(N, Q_0^2) = u_v^A(N, Q_0^2) + d_v^A(N, Q_0^2)$$
(15)

Here $a_s(=\alpha_s/4\pi)$ and a_0 represent the strong coupling constant at the scales of Q^2 and Q_0^2 respectively. $C_{3,NS}^{(m)}(N)$ refers to the nonsinglet Wilson coefficients in $O(a_s^m)$, which can be found in the cited reference [40]. The term " \hat{P}_m " also denotes the Mellin transforms of the (m + 1)-loop splitting functions.

The strong coupling constant a_s is of utmost significance in the present paper regarding the evolution of parton densities. At $N^m LO$, the scale dependence of a_s

is determined by

$$\frac{d a_s}{d \ln Q^2} = \beta_{N^m LO}(a_s) = -\sum_{k=0}^m a_s^{k+2} \beta_k .$$
 (16)

In complete N³LO-loop approximation and using the Λ -

parametrization, the running coupling is given by [41, 42]:

$$a_{s}(Q^{2}) = \frac{1}{\beta_{0}L_{\Lambda}} - \frac{1}{(\beta_{0}L_{\Lambda})^{2}} b_{1}\ln L_{\Lambda} + \frac{1}{(\beta_{0}L_{\Lambda})^{3}} \\ \left[b_{1}^{2}\left(\ln^{2}L_{\Lambda} - \ln L_{\Lambda} - 1\right) + b_{2}\right] + \frac{1}{(\beta_{0}L_{\Lambda})^{4}} \\ \left[b_{1}^{3}\left(-\ln^{3}L_{\Lambda} + \frac{5}{2}\ln^{2}L_{\Lambda} + 2\ln L_{\Lambda} - \frac{1}{2}\right) \\ - 3b_{1}b_{2}\ln L_{\Lambda} + \frac{b_{3}}{2}\right], \qquad (17)$$

where $L_{\Lambda} \equiv ln(Q^2/\Lambda^2)$, $b_k \equiv \beta_k/\beta_0$, and Λ is the QCD scale parameter. The first line of Eq. (17) includes the the NLO-loop coefficients, the second line is the N²LOloop and the third line denotes the N³LO-loop correction. Equation (17) solves the evolution equation (16) only up to higher orders in $1/L_{\Lambda}$. The functional form of $\alpha_s(Q^2)$, in N³LO-loop approximation and for 6 different values of Λ . To be able to compare with other measurements of Λ we adopt the matching of flavor thresholds at $Q^2 = m_c^2$ and $Q^2 = m_b^2$ with $m_c = 1.5$ GeV and $m_b = 4.5$ GeV as described in [43, 44].

III. JACOBI POLYNOMIALS AND THE PROCEDURE OF QCD FITS

One of the simplest and fastest methods for reconstructing the structure function from QCD predictions for its Mellin moments is through the expansion of Jacobi polynomials. The Jacobi polynomials are especially suitable for this purpose since they allow one to factor out an essential part of the x-dependence of structure function into the weight function [45].

According to this method, one can relate the xF_3 structure function with its Mellin moments

$$xF_{3}^{N_{max}}(x,Q^{2}) = x^{\beta}(1-x)^{\alpha} \sum_{n=0}^{N_{max}} \Theta_{n}^{\alpha,\beta}(x)$$
$$\sum_{j=0}^{n} c_{j}^{(n)}(\alpha,\beta)F_{3}(j+2,Q^{2}),$$
(18)

where N_{max} is the number of polynomials. Jacobi polynomials of order n [45], $\Theta_n^{\alpha,\beta}(x)$, satisfy the orthogonality condition with the weight function $w^{\alpha\beta} = x^{\beta}(1-x)^{\alpha}$

$$\int_0^1 dx \ w^{\alpha\beta} \Theta_k^{\alpha,\beta}(x) \Theta_l^{\alpha,\beta}(x) = \delta_{k,l} \ . \tag{19}$$

In the above, $c_j^{(n)}(\alpha,\beta)$ are the coefficients expressed through Γ -functions and satisfying the orthogonality relation in Eq. (19) and $F_3(j+2,Q^2)$ are the moments determined in the previous section. N_{max} , α and β have to be chosen so as to achieve the fastest convergence of the series on the right-hand side of Eq. (18) and to reconstruct F_2 with the required accuracy. In our analysis we

use $N_{max} = 9$, $\alpha = 3.0$ and $\beta = 0.5$. The same method has been applied to calculate the nonsinglet structure function xF_3 from their moments [9, 12, 15, 16] and for polarized structure function xg_1 [46–49]. Obviously the Q^2 -dependence of the polarized structure function is defined by the Q^2 -dependence of the moments. The evolution equations allow for the calculation of the Q^2 dependence of parton distributions provided at a certain reference point, Q_0^2 . These distributions are typically parameterized based on plausible theoretical assumptions regarding their behavior near the endpoints, x = 0 and x = 1. For the data utilized in the global analysis, most experiments combine various systematic errors into one effective error for each data point, along with the statistical error. Additionally, the fully correlated normalization error of the experiment is usually specified separately. Therefore, it is natural to adopt the following definition for the effective χ^2 [50].

$$\chi^{2}_{\text{global}} = \sum_{n} w_{n} \chi^{2}_{n} , \text{ (n labels the different experiments)}$$
$$\chi^{2}_{n} = \left(\frac{1 - \mathcal{N}_{n}}{\Delta \mathcal{N}_{n}}\right)^{2} + \sum_{i} \left(\frac{\mathcal{N}_{n} x F^{data}_{3,i} - x F^{theor}_{3,i}}{\mathcal{N}_{n} \Delta x F^{data}_{3,i}}\right)^{2} . \tag{20}$$

For the n^{th} experiment, $xF_{3,i}^{data}$, $\Delta xF_{3,i}^{data}$, and $xF_{3,i}^{theor}$ denote the data value, measurement uncertainty (statistical and systematic combined) and theoretical value for the i^{th} data point. $\Delta \mathcal{N}_n$ is the experimental normalization uncertainty and \mathcal{N}_n is an overall normalization factor for the data of experiment n. The factor w_n is a possible weighting factor (with default value 1). However, we allowed for a relative normalization shift \mathcal{N}_n between the different data sets within the normalization uncertainties $\Delta \mathcal{N}_n$ quoted by the experiments.

 $\Delta \mathcal{N}_n$ quoted by the experiments. Now the sums in χ^2_{global} run over all data sets and in each data set over all data points. The minimization of the above χ^2 value to determine the best parametrization of the unpolarized parton distributions is done using the program MINUIT [51].

IV. FIT RESULTS

The data for the charged-current structure functions $xF_3(x, Q^2)$ used in our analysis are listed in Table. I. The x and Q^2 ranges, the number of data points and the related references are also listed in this table. The CCFR [25] and NuTeV [26] collaborations at Fermilab conduct neutrino deep inelastic scattering experiments using an iron target, which are subsequently adjusted to account for an isoscalar target. They cover much of the same kinematic range of momentum fraction x, but CCFR covers slightly higher Q^2 . At high values of x, the predictions are mainly determined by the valence up quark distribution, which is very well constrained by the charged-current DIS structure function data. We also

Experiment	x	Q^2	Number of data points	Reference
CCFR	$0.0075 \le x \le 0.75$	$1.3 \le Q^2 \le 125.9$	116	[25]
NuTeV	$0.015 \le x \le 0.75$	$3.162 \le Q^2 \le 50.118$	64	[26]
CHORUS	$0.02 \le x \le 0.65$	$2.052 \le Q^2 \le 81.55$	50	[27]

Table I: Published data points for charged-current structure functions $xF_3(x, Q^2)$ used in the present global fit. The x and Q^2 ranges, the number of data points and the related references are also listed.

include recent data from the CHORUS [27] collaboration, which are taken from a lead target and cover a similar range in x compared with CCFR. The NuTeV data seem to be more precise. In practice, we find the high-x NuTeV and CHORUS data very difficult to fit, leading to higher values of χ^2 .

The accuracy of this result stems from the obtained $\chi^2/d.o.f$ value by fitting the initial valence PDFs at $Q_0^2 = 1 \ GeV^2$, as detailed in Table II. The world average value for $\alpha_s(M_z^2) = 0.1179 \pm 8.5 \times 10^{-6}$, as reported in Ref. [58] is in good agreement with the reported ones in Table II. In Figs. 3, 4 and 5, valence PDFs for different approaches at $Q^2 = 1 \ \text{GeV}^2$ and $Q^2 = 10 \ \text{GeV}^2$ have been depicted in NLO, NNLO, and NNNLO approximations. As can be seen, the agreement with the results of MSHT20 [53] improves notably at NNLO. The results of the analysis by MSTH23 [54] in N³LO are very close to those of AMGA24 at $Q^2 = 1 \ \text{GeV}^2$.

Additionally, the results for the xF_3 proton structure function in NLO, NNLO and NNNLO approximations are presented in Fig. 6. The comparison to experimental data from the CCFR collaboration [25], which provides direct measurements of the xF_3 structure function, is also shown. The agreement between the theoretical predictions and the experimental data highlights the reliability of the current theoretical framework.

Furthermore, in Figs. 7 and 8, the xF_3 structure functions at NLO, NNLO, and NNNLO are compared with experimental data from the NuTeV [26] and CHORUS [27] collaborations at various values of Q^2 . These comparisons further validate the theoretical predictions and provide additional constraints on the PDFs. In Ref. [3], the χ^2 for NNLO was reported as 1.482, while in the current model, we obtain 1.4691, indicating that the choice of Gegenbauer polynomial is appropriate. We have done calculations up to NNNLO order accurately. As can be observed, our analysis derives the exact solution by incorporating the splitting function as per Ref. [36] and the Wilson coefficient following Ref. [40] within the NNNLO approximation, unlike Ref. [16] which employed the Padé approximation method for the xF_3 structure function at NNNLO. We have established that the Gegenbauer polynomial is as effective as the Chebyshev polynomial, corroborated by findings from the MSHT20 and MSTH23 groups in Refs. [53] and [54], respectively. Our analysis successfully computes the non-singlet component of the structure function using a minimal dataset and a streamlined approach, achieving congruence with other established models.



Figure 3: xu_v and xd_v in NLO and N²LO approximation in Mellin model compare with MSHT20 [53] and CT18 [52]



Figure 4: xu_v and xd_v in N³LO approximation in Mellin model compare with MSHT23 [54]

V. THE GROSS-LLEWELLYN SMITH (GLS) SUM RULE

Another intriguing issue revolves around the extraction of the value of the Gross-Llewellyn Smith (GLS) sum rule. The GLS sum rule is a crucial property in the context of deep inelastic neutrino-nucleon scattering. In the framework of the quark-parton model, the GLS sum rule, which is associated with the xF_3 structure function, is expressed as [55].

The GLS sum rule relates the integral of the $xF_3(x, Q^2)$

Table II: Best fit parameters and uncertainties of the Mellin fits at NLO , N²LO and N³LO at the initial scale $Q_0^2 = 1.0 \text{ GeV}^2$.

		Mellin	
	NLO	$N^{2}LO$	N ³ LO
a_u^1	-0.20077 ± 0.0013422	0.026755 ± 0.00332961	-0.16898 ± 0.022263
a_u^2	0.030599 ± 0.000367904	$-0.0195933 \pm 0.00170209$	0.0235584 ± 0.00059855
a_u^3	$-0.0033036 \pm 0.00010995$	$0.003302035 \pm 0.000536726$	$-0.0020921 \pm 0.00017266$
α_{uv}	0.63134 ± 0.008237101	0.782549 ± 0.00586995	0.78679 ± 0.016633
β_{uv}	4.05839 ± 0.055065	2.962457 ± 0.0272975	4.24605 ± 0.057478
β_{d_v}	-0.0075767 ± 0.16530	0.00262071 ± 0.082775	0.00105091 ± 0.17327
$\alpha_s(Q_0^2)$	0.49941 ± 0.017535	0.45056 ± 0.007143	0.41758 ± 0.01003
$\alpha_s(M_z^2)$	0.12101 ± 0.000551	0.11958 ± 0.007143	0.118011 ± 0.000505
$\chi^2/d.o.f$	368.8118/223 = 1.6538	327.6235/223 = 1.4691	310.0293/223 = 1.39026



Figure 5: xu_v and xd_v at $Q^2 = 10GeV^2$ in NLO, N²LO and N³LO approximation in Mellin model compare with MSHT20 [53] and MSHT23 [54]

structure function over the entire x:

$$GLS(Q^2) = \frac{1}{2} \int_0^1 \frac{x F_3^{\bar{\nu}p+\nu p}(x,Q^2)}{x} dx. \qquad (21)$$

Its experimental verification provides valuable constraints on the parameters of the electroweak theory and is crucial for understanding the interplay between weak and strong interactions in the nucleon. By accurately ex-



Figure 6: xF_3 in NLO , N²LO and N³LO approximation in Mellin model compare with CCFR data [25]

tracting the value of the GLS sum rule from experimental data and comparing it with theoretical predictions, we can gain insights into the quark and parton distributions inside the nucleon and shed light on the physics beyond the Standard Model.

In the work of Ref. [56], authors reported the following result for the measurement of the GLS sum rule at the scale $|Q^2| = 3 \text{ GeV}^2$,

GLS
$$(|Q^2| = 3 \text{ GeV}^2) = 2.5 \pm 0.018 \text{ (stat.)} \pm 0.078 \text{ (syst.)}.$$
(22)

The value of the GLS sum rule at the scale $|Q^2| = 8$ GeV² is reported as 2.62 ± 0.15 in Ref. [57]. In our work, we obtained GLS, $(|Q^2| = 8, \text{GeV}^2) = 2.46591 \pm 0.06289$



Figure 7: xF_3 at NLO, N²LO and N³LO approximation in Mellin model compare with NuTeV data [26].



Figure 8: xF_3 at NLO , N²LO and N³LO approximation in Mellin model compare with CHORUS data [27].

for the NLO analysis GLS, $(|Q^2| = 8, \text{GeV}^2) = 2.46271 \pm$

0.04021 and for N²LO analysis ,GLS, $(|Q^2| = 8, \text{GeV}^2) = 2.32743 \pm 0.03761$ and for N³LO analysis in Mellin space which are in good agreement with the results obtained by mentioned research groups.

VI. SUMMARY AND CONCLUSIONS

In this paper, we present an analysis of the valence quark distribution functions in the proton using Gegenbauer polynomials for their parameterization. These polynomials have many advantages in QCD analysis; they provide a flexible framework for expanding functions, allowing the approximation of various shapes and behaviors in the PDFs, and facilitating the fitting of experimental data and extraction of relevant physical quantities. Moreover, Gegenbauer polynomials enable the systematic incorporation of higher-order NNNLO corrections, thereby enhancing the precision of theoretical predictions in QCD analyses. In our analysis, we employ precise splitting functions and Wilson coefficients, eschewing the use of Padé approximation, at the Next-to-Next-to-Next-to-Leading Order. In fact we demonstrated that using the Gegenbauer polynomial for parameterization is a fast, precise, and direct way to calculate the final structure function with high accuracy. The comparison with the MSTH23 and MSTH20 results, which used the Chebyshev polynomials, shows the advantage of the Gegenbauer polynomials and provides a more precise determination of xF_3 in the kinematic region of interest without complicated calculations.

Through careful analysis, the figures presented in our study show a convincing correspondence with both established models and empirical data, attesting to the robustness and reliability of our methodology. This concordance underscores the validity of our findings and supports the wider applicability of our approach within the field of particle physics.

Appendix A: FORTRAN PACKAGE OF OUR NLO , NNLO and NNNLO PDFS

A FORTRAN package containing our unpolarized PDFs a at NLO, NNLO and NNNLO approximation as well as the unpolarized structure functions $xF_3(x, Q^2)$ can be obtained via Email from the authors upon request. This package includes an example program to illustrate the use of the routines.

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- S. M. Moosavi Nejad, H. Khanpour, S. Atashbar Tehrani and M. Mahdavi, Phys. Rev. C 94, no.4, 045201 (2016).
- [2] A. N. Khorramian and S. Atashbar Tehrani, JHEP 03, 051 (2007).
- [3] A. G. Tooran, A. Khorramian and H. Abdolmaleki, Phys. Rev. C 99, no.3, 035207 (2019).
- [4] A. V. Sidorov and O. P. Solovtsova, "The QCD analysis of xF_3 structure function based on the analytic approach,"
- [5] A. V. Sidorov and O. P. Solovtsova, Mod. Phys. Lett. A 29, no.36, 1450194 (2014).
- [6] N. M. Nath, N. Baruah and J. K. Sarma, [arXiv:1101.2838 [hep-ph]].
- [7] A. L. Kataev, G. Parente and A. V. Sidorov, Nucl. Phys. B Proc. Suppl. 116, 105-109 (2003).
- [8] A. L. Kataev, G. Parente and A. V. Sidorov, J. Phys. G 29, 1985-1988 (2003).
- [9] A. L. Kataev, G. Parente and A. V. Sidorov, Phys. Part. Nucl. 34, 20-46 (2003). [erratum: Phys. Part. Nucl. 38, no.6, 827-827 (2007)]
- [10] J. Santiago and F. J. Yndurain, Nucl. Phys. B 611, 447-466 (2001).
- [11] S. I. Alekhin and A. L. Kataev, Nucl. Phys. A 666, 179-183 (2000).
- [12] A. L. Kataev, G. Parente and A. V. Sidorov, Nucl. Phys. B 573, 405-433 (2000).
- [13] A. V. Sidorov and M. V. Tokarev, Nuovo Cim. A **110**, 1401-1409 (1997).
- [14] A. L. Kataev, A. V. Kotikov,
 G. Parente and A. V. Sidorov,
 Nucl. Phys. B Proc. Suppl. 64, 138-142 (1998).
- [15] A. L. Kataev, A. V. Kotikov, G. Parente and A. V. Sidorov, Phys. Lett. B 417, 374-384 (1998).
- [16] A. L. Kataev, G. Parente and A. V. Sidorov, [arXiv:hep-ph/9809500 [hep-ph]].
- [17] A. V. Sidorov, JINR Rapid Commun. 80, 11 (1996)
- [18] A. L. Kataev, A. V. Kotikov, G. Parente and A. V. Sidorov, Phys. Lett. B 388,179-187 (1996).
 [10] A. L. Kataev, A. V. Kotikov, G. Parente and A. V. Sidorov, Phys. Lett. B 388,179-187 (1996).
- [19] A. L. Kataev and A. V. Sidorov, Phys. Lett. B **331**,179-186 (1994).
- [20] B. R. Martin and G. Shaw, Z. Phys. C **33**,99 (1986).
- [21] A. Behring, J. Blümlein, A. De Freitas, A. Hasselhuhn, A. von Manteuffel and C. Schneider, Phys. Rev. D 92,no.11, 114005 (2015).
- [22] S. Atashbar Tehrani, Phys. Rev. C 86,064301 (2012).
- [23] H. Khanpour and S. Atashbar Tehrani, Phys. Rev. D 93,no.1, 014026 (2016).
- [24] M. Gluck, E. Reya and A. Vogt, Z. Phys. C 48,471-482 (1990).
- [25] W. G. Seligman, C. G. Arroyo, L. de Barbaro, P. de Barbaro, A. O. Bazarko, R. H. Bernstein, A. Bodek, T. Bolton, H. S. Budd and J. Conrad, *et al.* Phys. Rev. Lett. **79**,1213-1216 (1997).
- [26] M. Tzanov itet al. [NuTeV], Phys. Rev. D 74, 012008 (2006).
- [27] G. Onengut itet al. [CHORUS], Phys. Lett. B 632, 65-75 (2006).
- [28] J. Blumlein, H. Bottcher and A. Guffanti, Nucl. Phys. B 774, 182-207 (2007).
- [29] J. Blümlein and M. Saragnese, Phys. Lett. B 820, 136589 (2021).

- [30] S. Moch and J. A. M. Vermaseren, Nucl. Phys. B 573, 853-907 (2000).
- [31] W. L. van Neerven and A. Vogt, Nucl. Phys. B 568, 263-286 (2000).
- [32] E. G. Floratos, C. Kounnas and R. Lacaze, Nucl. Phys. B **192**, 417-462 (1981).
- [33] S. Moch, J. A. M. Vermaseren and A. Vogt, Nucl. Phys. B 688, 101-134 (2004).
- [34] A. Vogt, S. Moch and J. A. M. Vermaseren, Nucl. Phys. B 691, 129-181 (2004).
- [35] P. Jimenez Delgado, [arXiv:0902.3947 [hep-ph]].
- [36] S. Moch, B. Ruijl, T. Ueda, J. A. M. Vermaseren and A. Vogt, JHEP 10, 041 (2017).
- [37] O. V. Tarasov, A. A. Vladimirov and A. Y. Zharkov, Phys. Lett. B 93, 429-432 (1980).
- [38] S. A. Larin and J. A. M. Vermaseren, Phys. Lett. B 303, 334-336 (1993).
- [39] J. Blümlein, P. Marquard, C. Schneider and K. Schönwald, Nucl. Phys. B 971, 115542 (2021).
- [40] J. Blümlein, P. Marquard, C. Schneider and K. Schönwald, JHEP 11, 156 (2022).
- [41] A. Vogt, Comput. Phys. Commun. 170, 65-92 (2005).
- [42] K. G. Chetyrkin, B. A. Kniehl and M. Steinhauser, Phys. Rev. Lett. **79**, 2184-2187 (1997).
- [43] W. A. Bardeen, A. J. Buras, D. W. Duke and T. Muta, Phys. Rev. D 18, 3998 (1978).
- [44] K. Abe et al. [E143 Collaboration], Phys. Lett. B 452, 194 (1999).
- [45] G. Parisi and N. Sourlas, Nucl. Phys. B **151**, 421 (1979).
- [46] H. Nematollahi, A. Mirjalili and S. Atashbar Tehrani, Phys. Rev. D 107, no.5, 054033 (2023).
- [47] A. Mirjalili and S. Tehrani Atashbar, Phys. Rev. D 105, no.7, 074023 (2022).
- [48] H. Nematollahi, P. Abolhadi, S. Atashbar, A. Mirjalili and M. M. Yazdanpanah, Eur. Phys. J. C 81, no.1, 18 (2021).
- [49] M. Salajegheh, S. M. Moosavi Nejad, M. Nejad, H. Khanpour and S. Atashbar Tehrani, Phys. Rev. C 97, no.5, 055201 (2018).
- [50] D. Stump, J. Pumplin, R. Brock, D. Casey, J. Huston, J. Kalk, H. L. Lai and W. K. Tung, Phys. Rev. D 65, 014012 (2002).
- [51] F. James and M. Roos, Comput. Phys. Commun. 10, 343-367 (1975).
- [52] T. J. Hou, J. Gao, T. J. Hobbs, K. Xie, S. Dulat, M. Guzzi, J. Huston, P. Nadolsky, J. Pumplin and C. Schmidt, *et al.* Phys. Rev. D **103**, no.1, 014013 (2021).
- [53] S. Bailey, T. Cridge, L. A. Harland-Lang, A. D. Martin and R. S. Thorne, the LHC," Eur. Phys. J. C 81, no.4, 341 (2021).
- [54] J. McGowan, T. Cridge, L. A. Harland-Lang and R. S. Thorne, Eur. Phys. J. C 83, no.3, 185 (2023). [erratum: Eur. Phys. J. C 83, no.4, 302 (2023)]
- [55] D. J. Gross and C. H. Llewellyn Smith, Nucl. Phys. B 14, 337-347 (1969).
- [56] W. C. Leung, P. Z. Quintas, S. R. Mishra, F. J. Sciulli, C. Arroyo, K. T. Bachmann, R. E. Blair, C. Foudas, B. J. King and W. C. Lefmann, *et al.* Phys. Lett. B **317**, 655-659 (1993).

- [57] J. T. Londergan and A. Phys. Rev. D 82, 113001 (2010).
 [58] P. A. Zyla *et al.* [Particle PTEP 2020, no.8, 083C01 (2020). W.Thomas,
- Data Group],